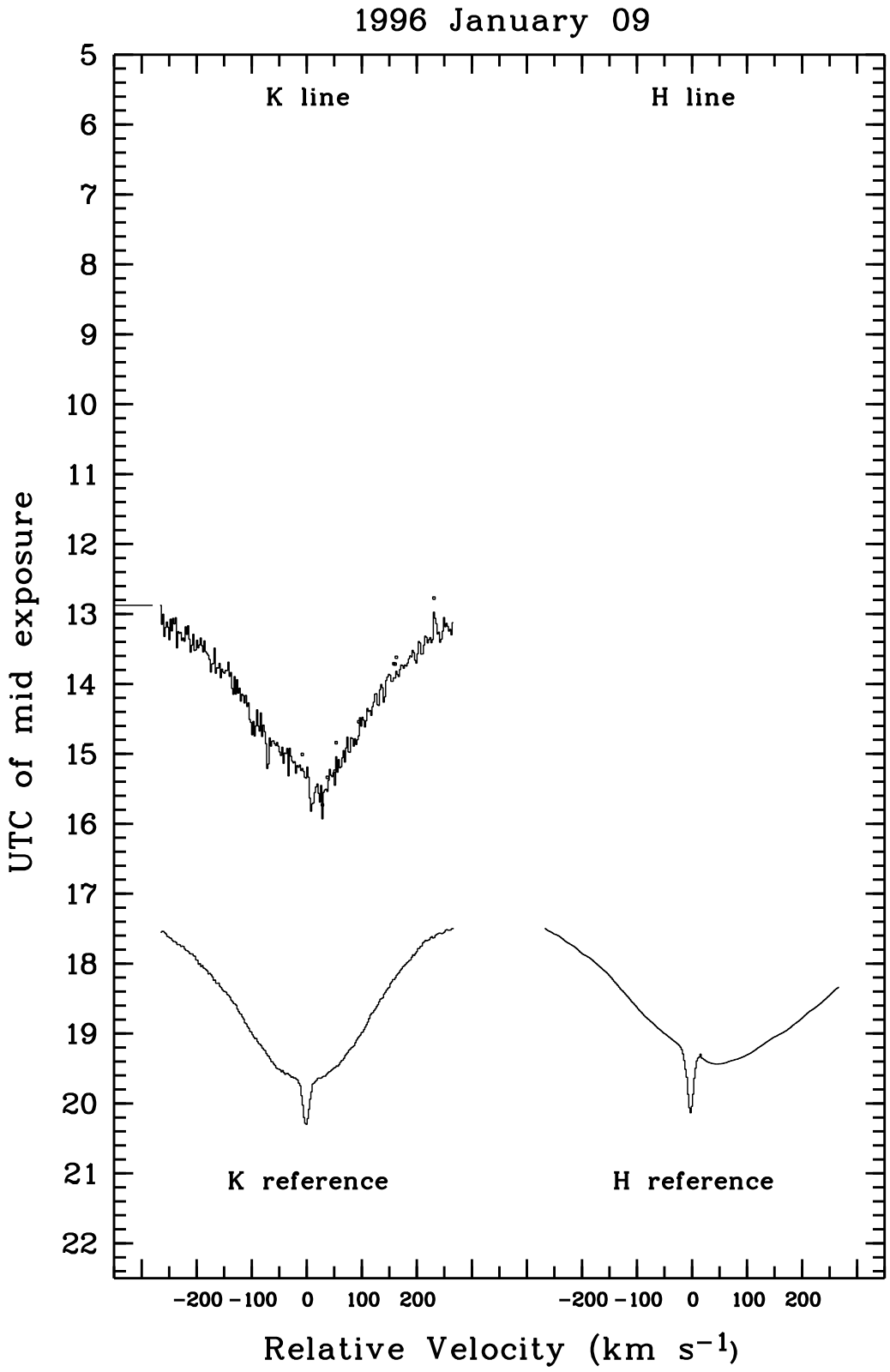

Figure 41: Observations of β Pictoris on 1996 January 9 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

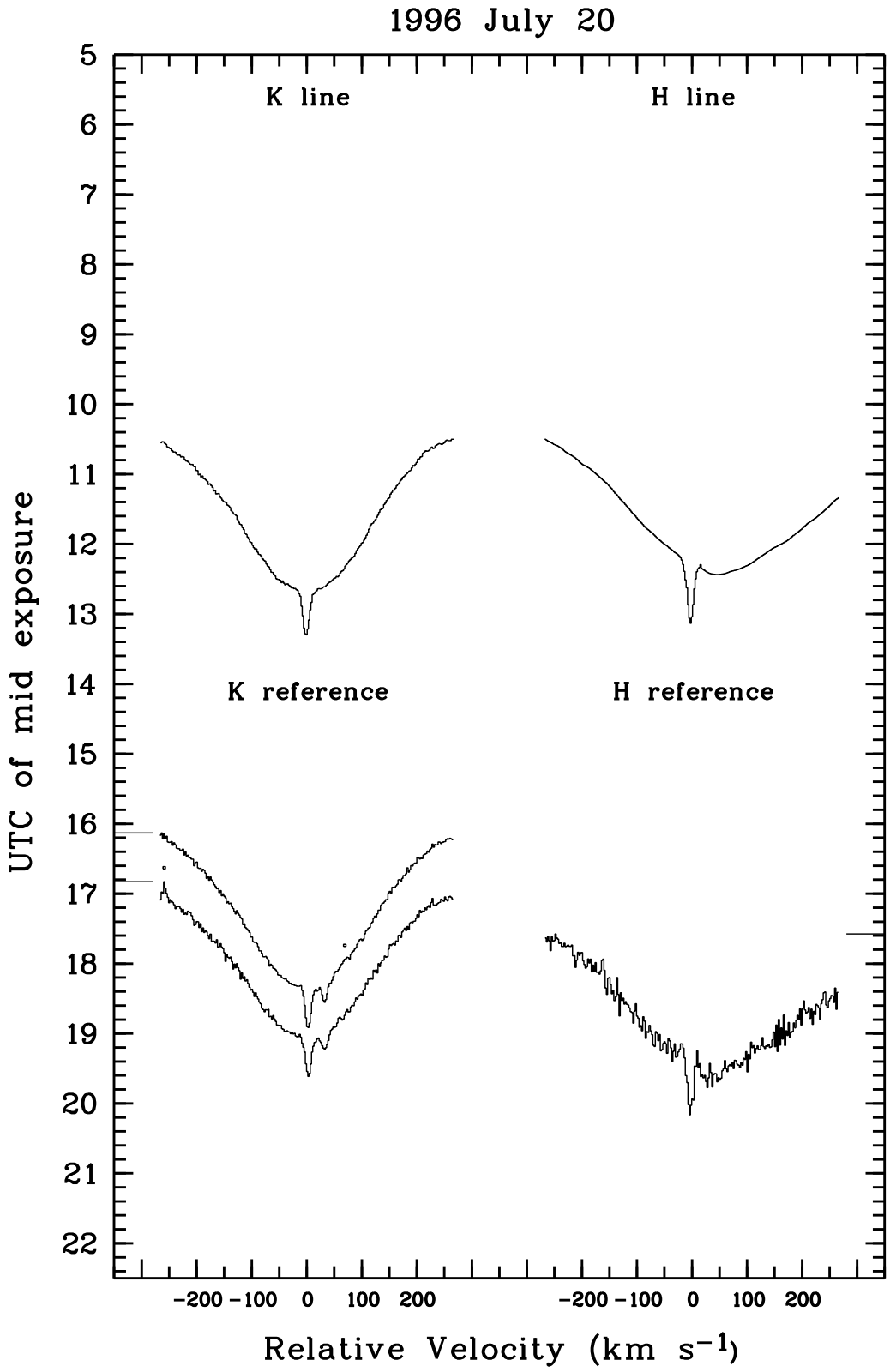

Figure 42: Observations of β Pictoris on 1996 July 20 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

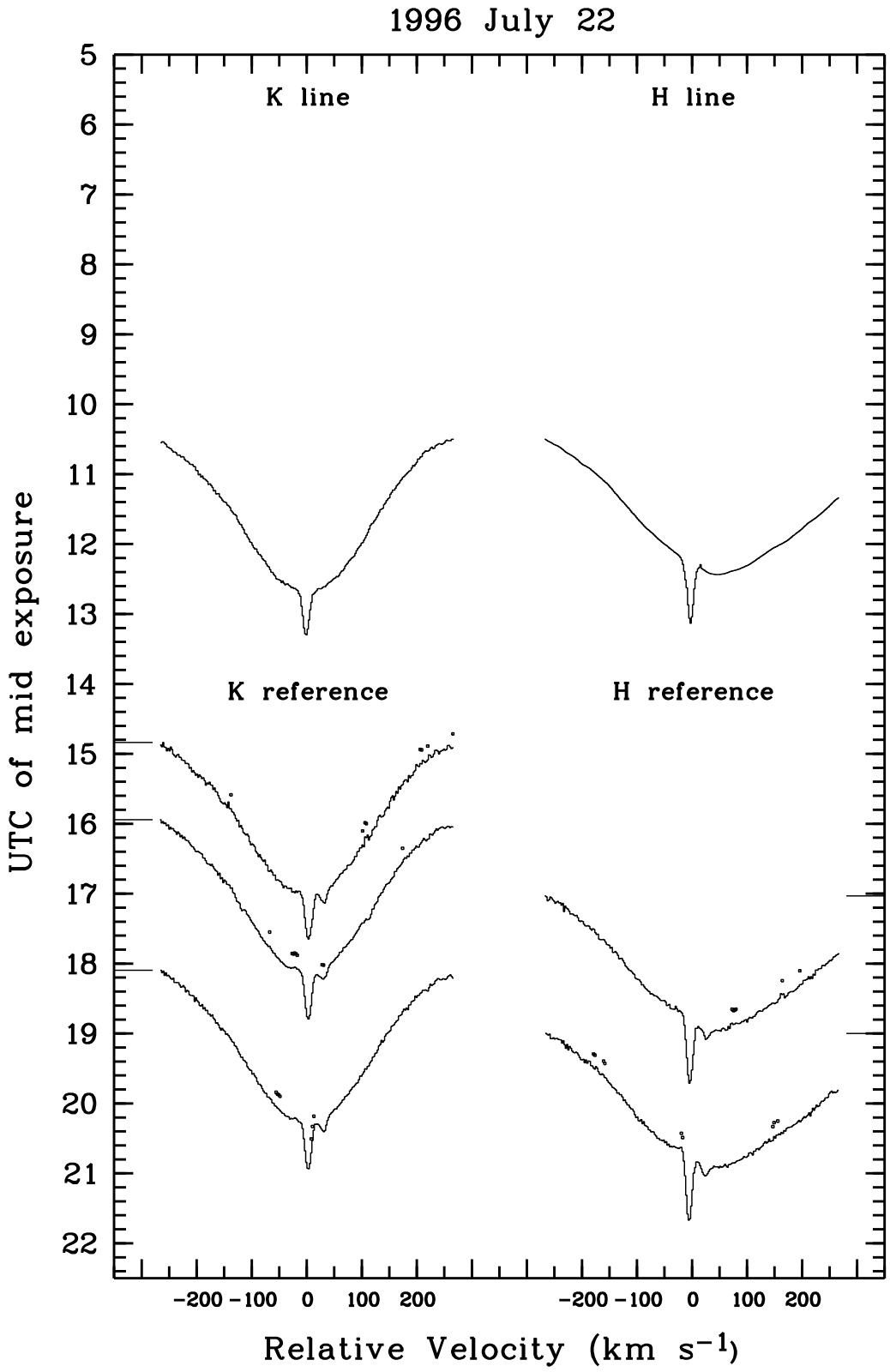

Figure 43: Observations of β Pictoris on 1996 July 22 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

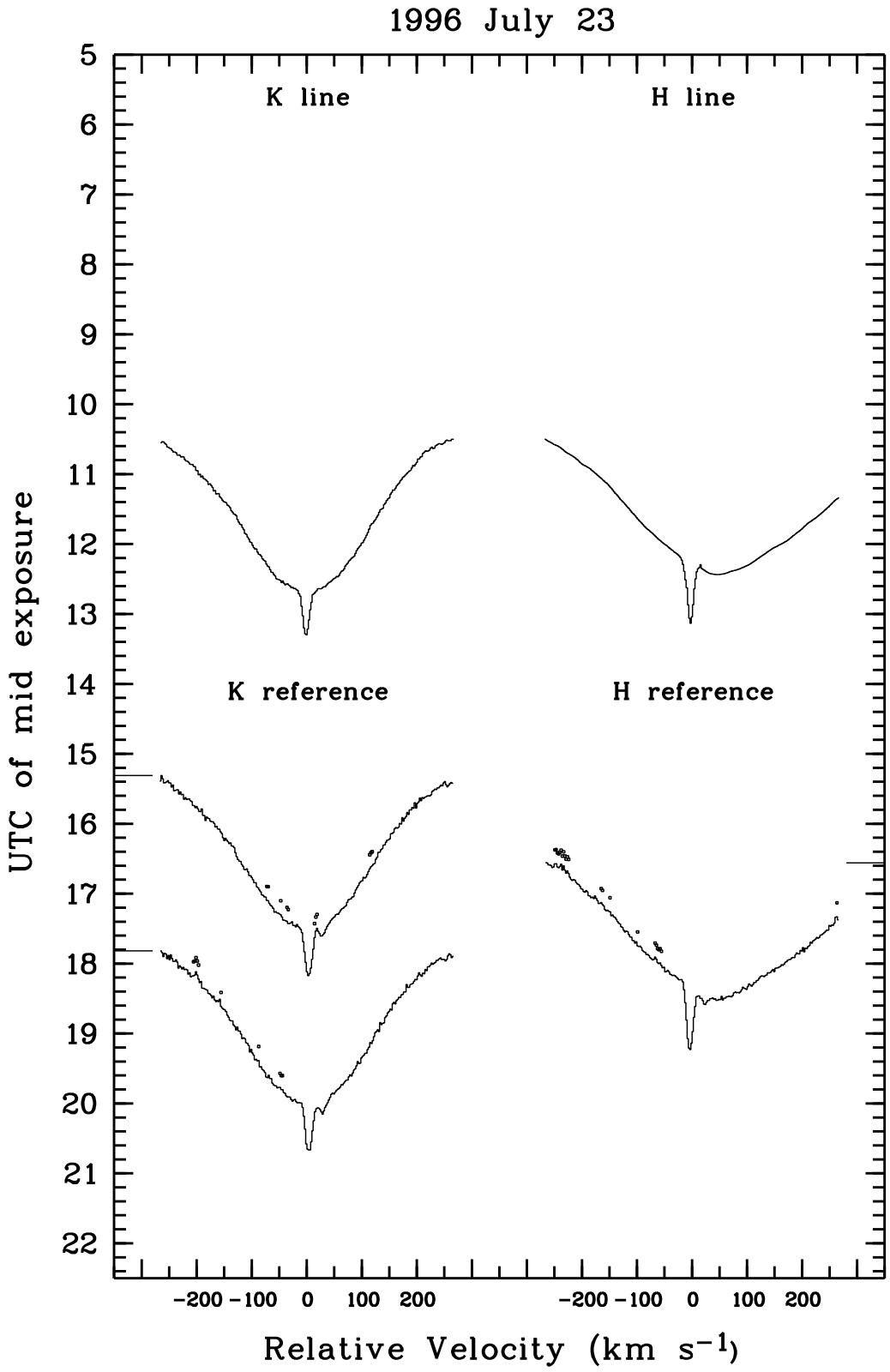

Figure 44: Observations of β Pictoris on 1996 July 23 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

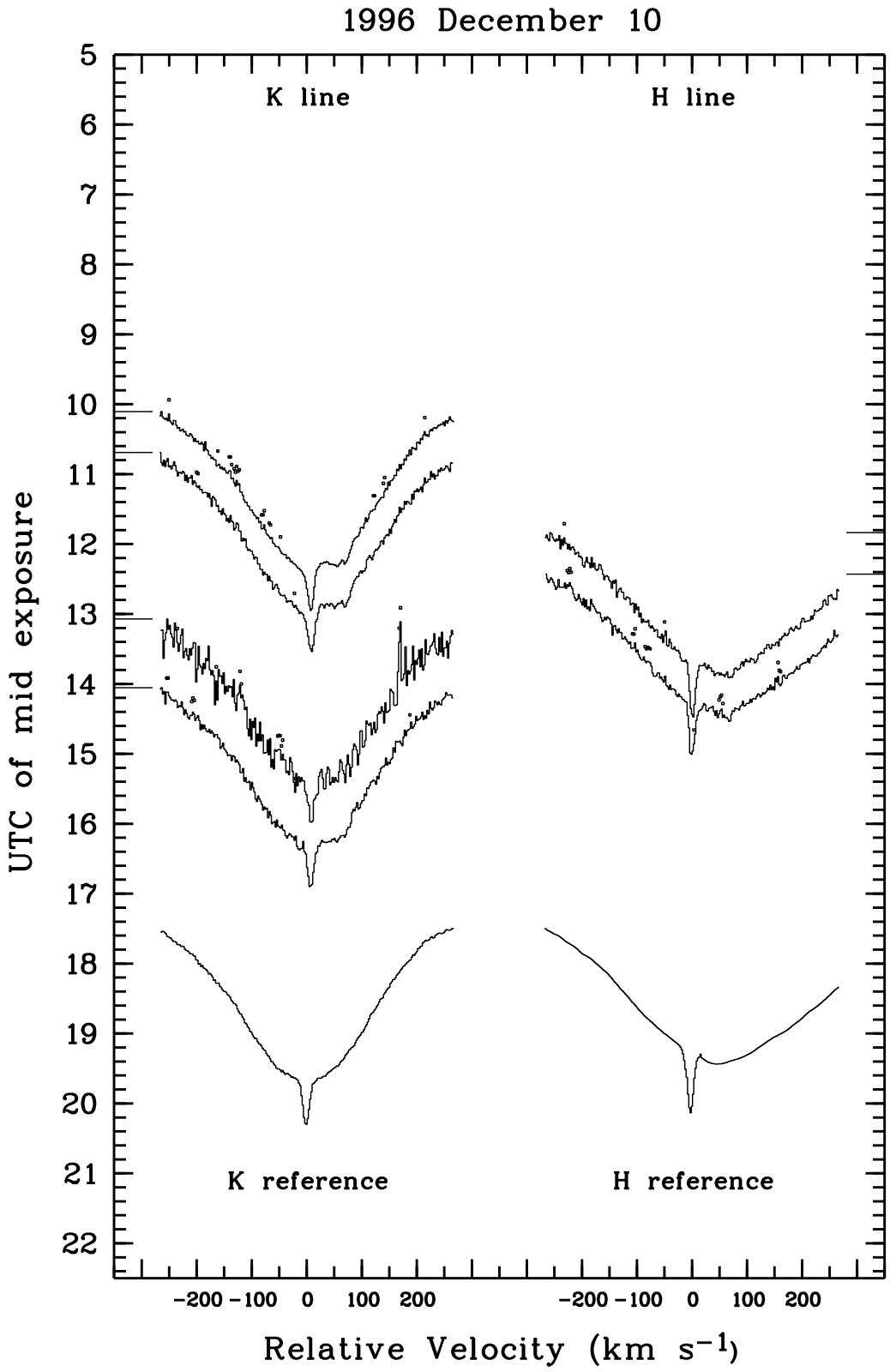

Figure 45: Observations of β Pictoris on 1996 December 10 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

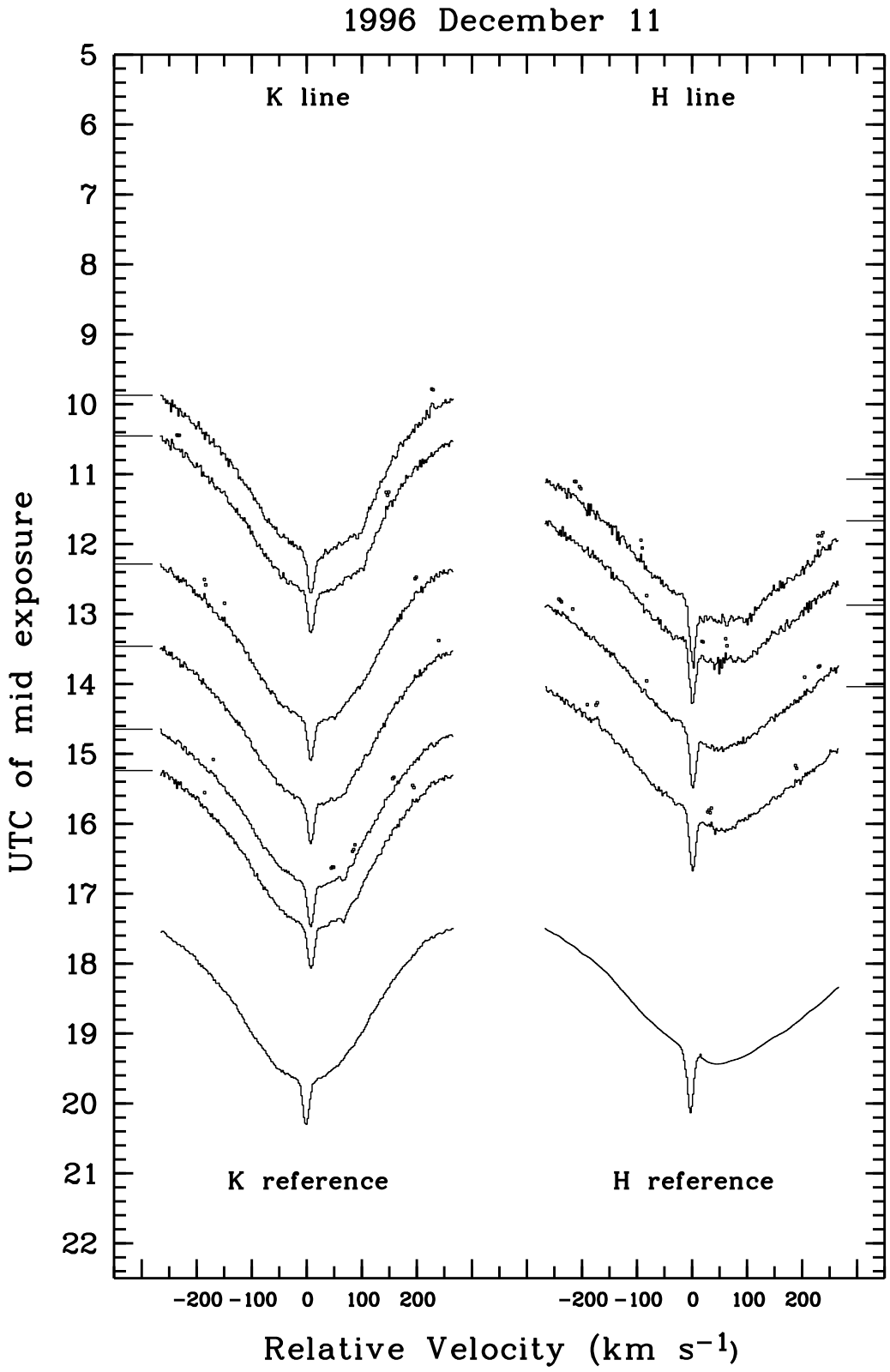

Figure 46: Observations of β Pictoris on 1996 December 11 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

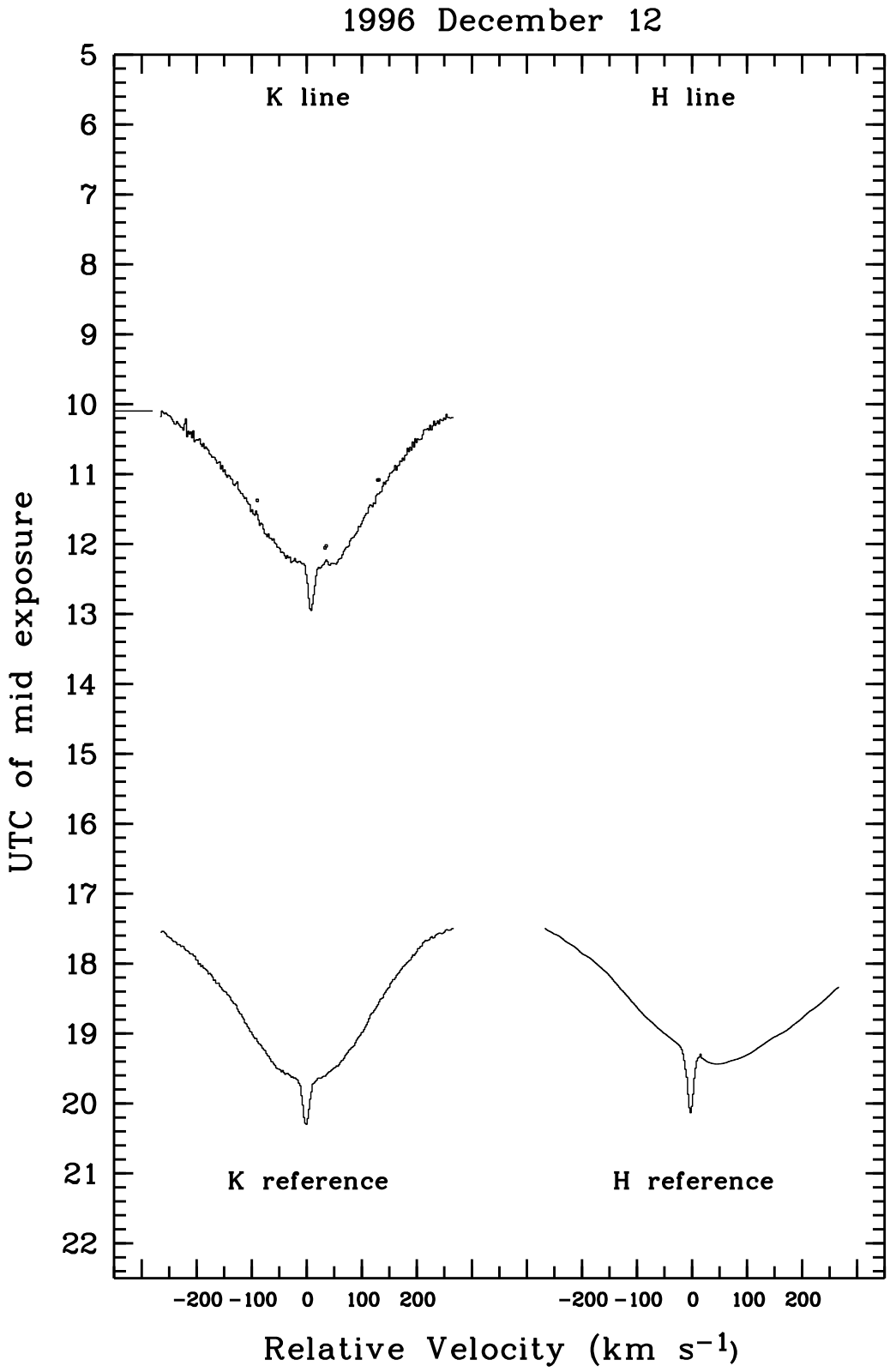

Figure 47: Observations of β Pictoris on 1996 December 12 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

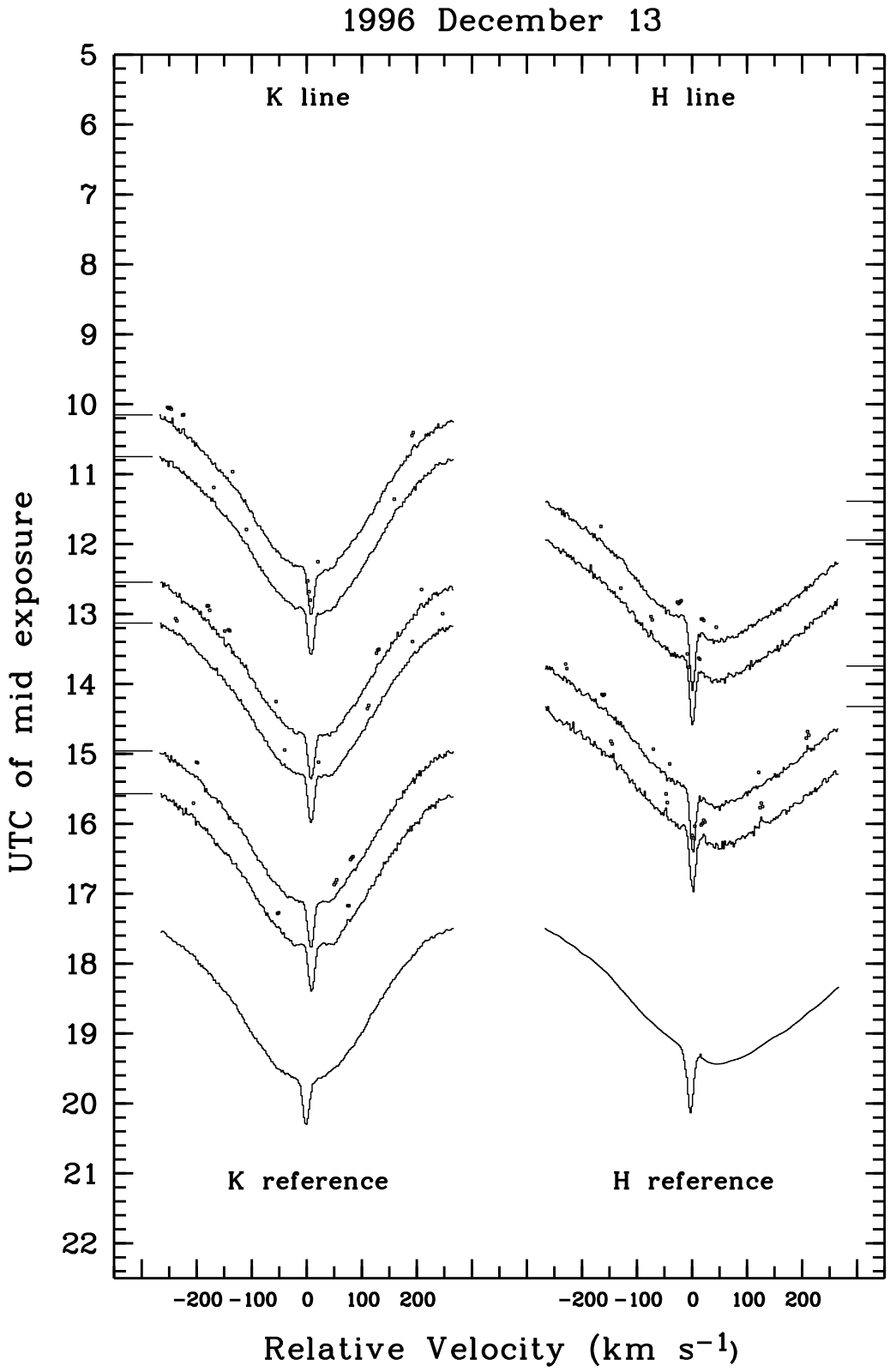

Figure 48: Observations of β Pictoris on 1996 December 13 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

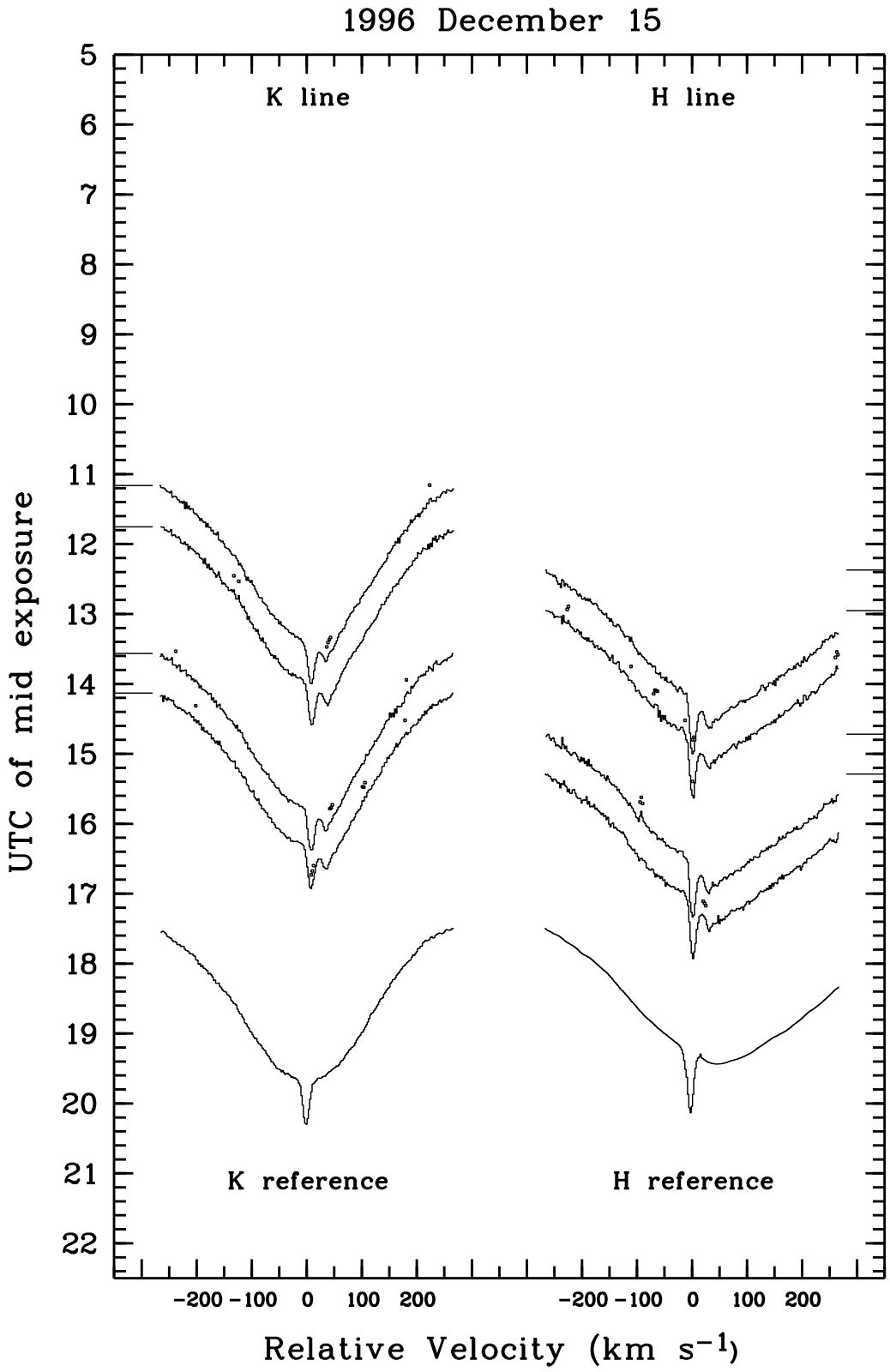

Figure 49: Observations of β Pictoris on 1996 December 15 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

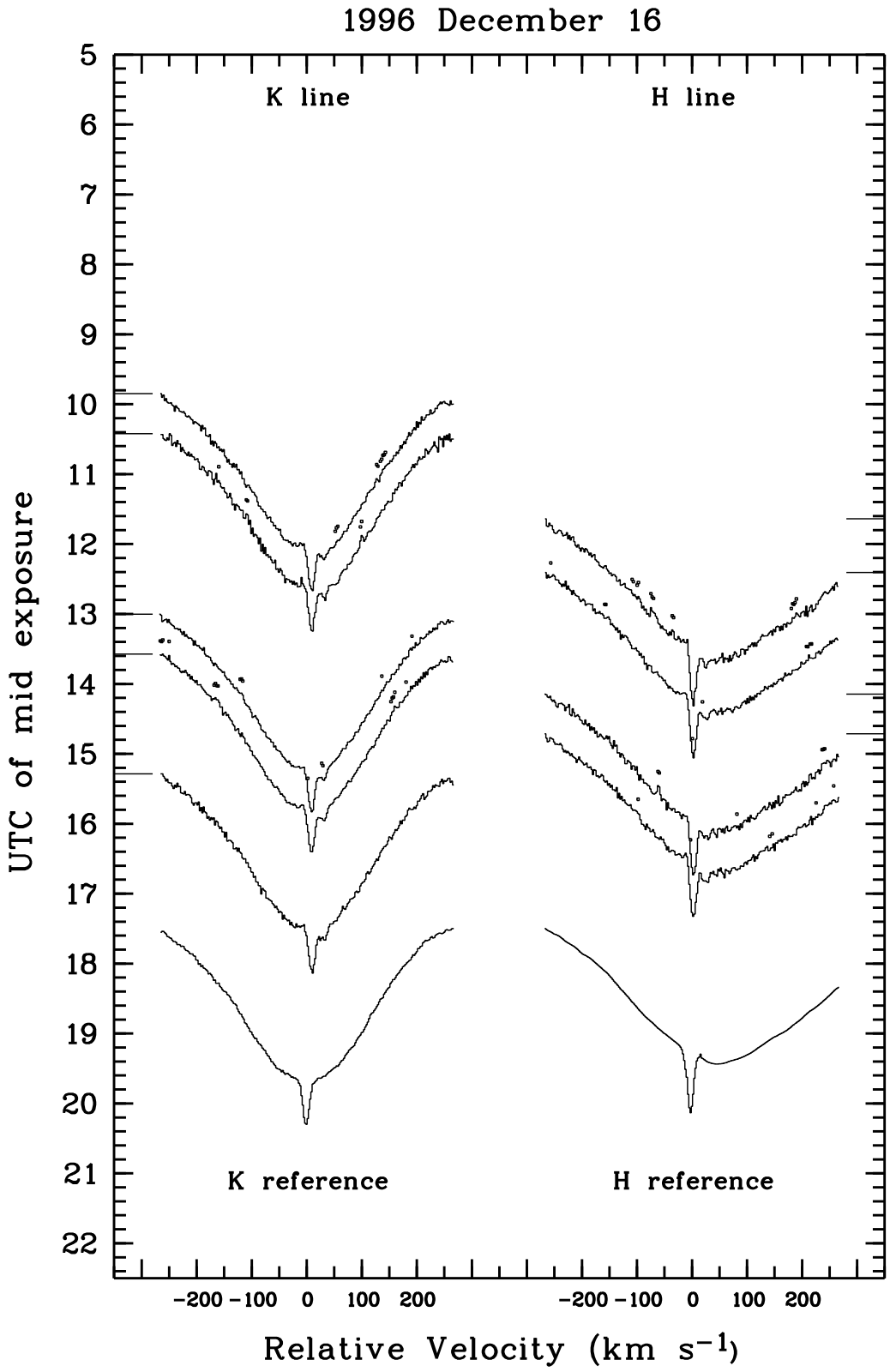

Figure 50: Observations of β Pictoris on 1996 December 16 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

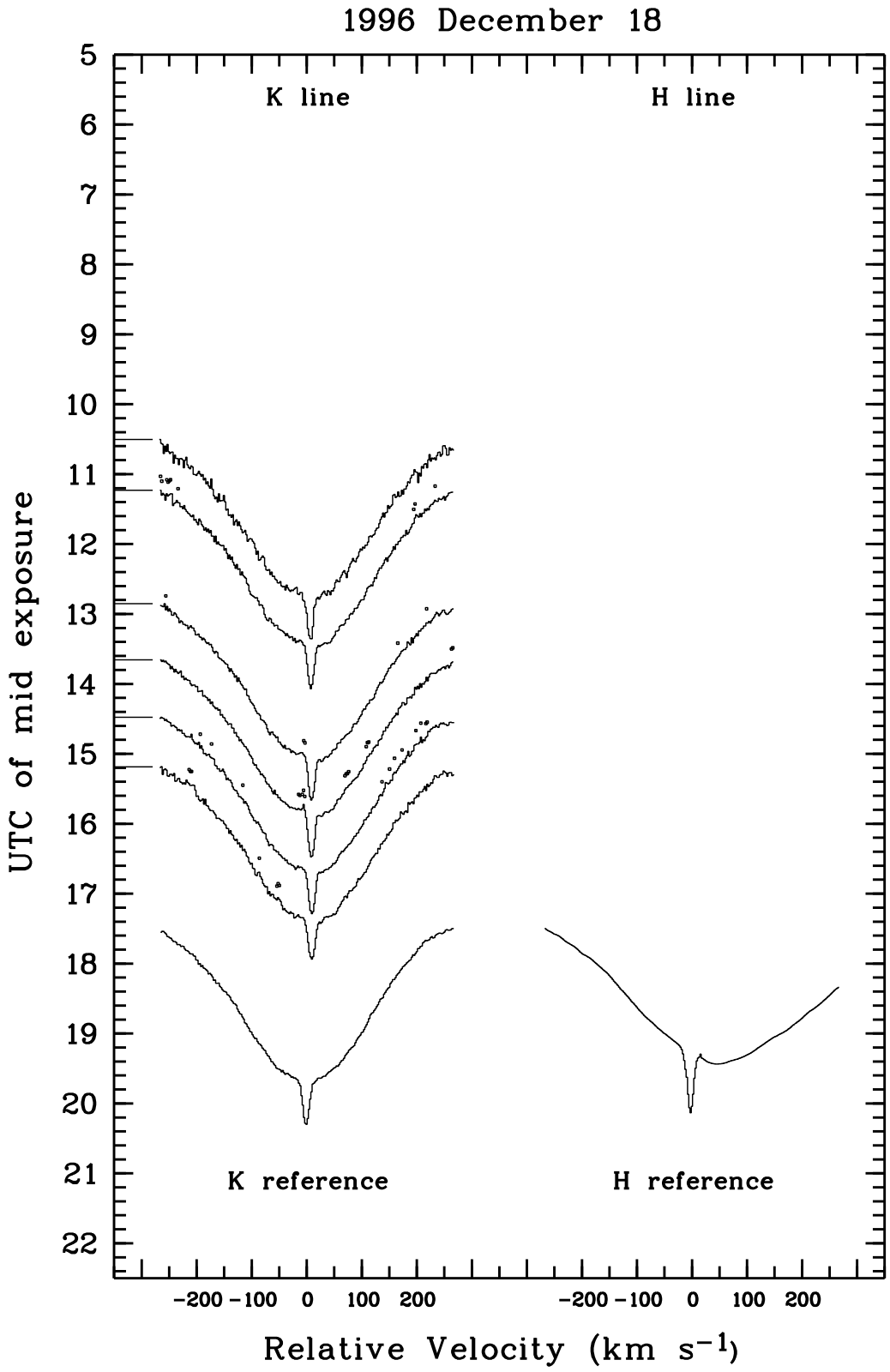

Figure 51: Observations of β Pictoris on 1996 December 18 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

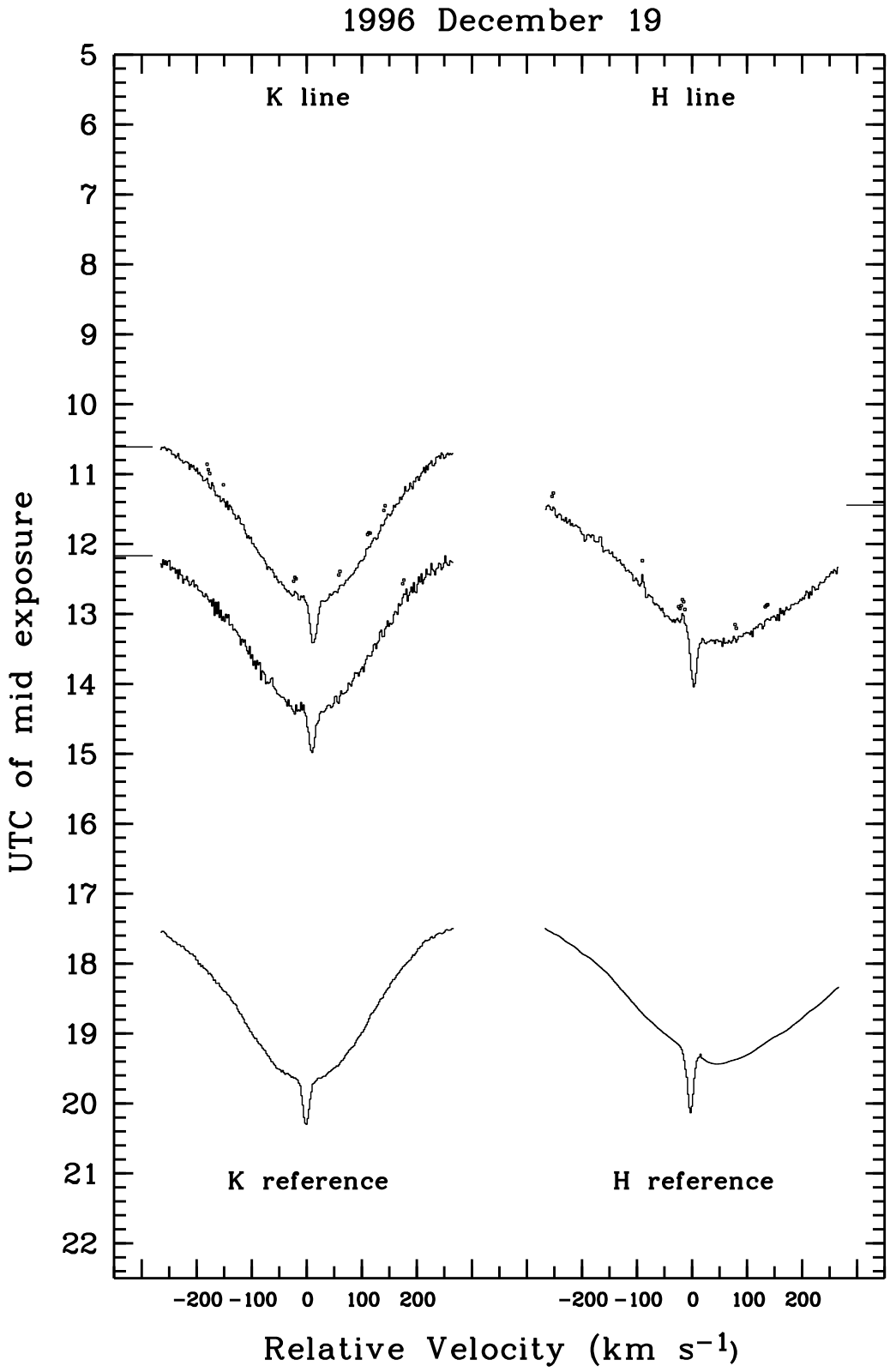

Figure 52: Observations of β Pictoris on 1996 December 19 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

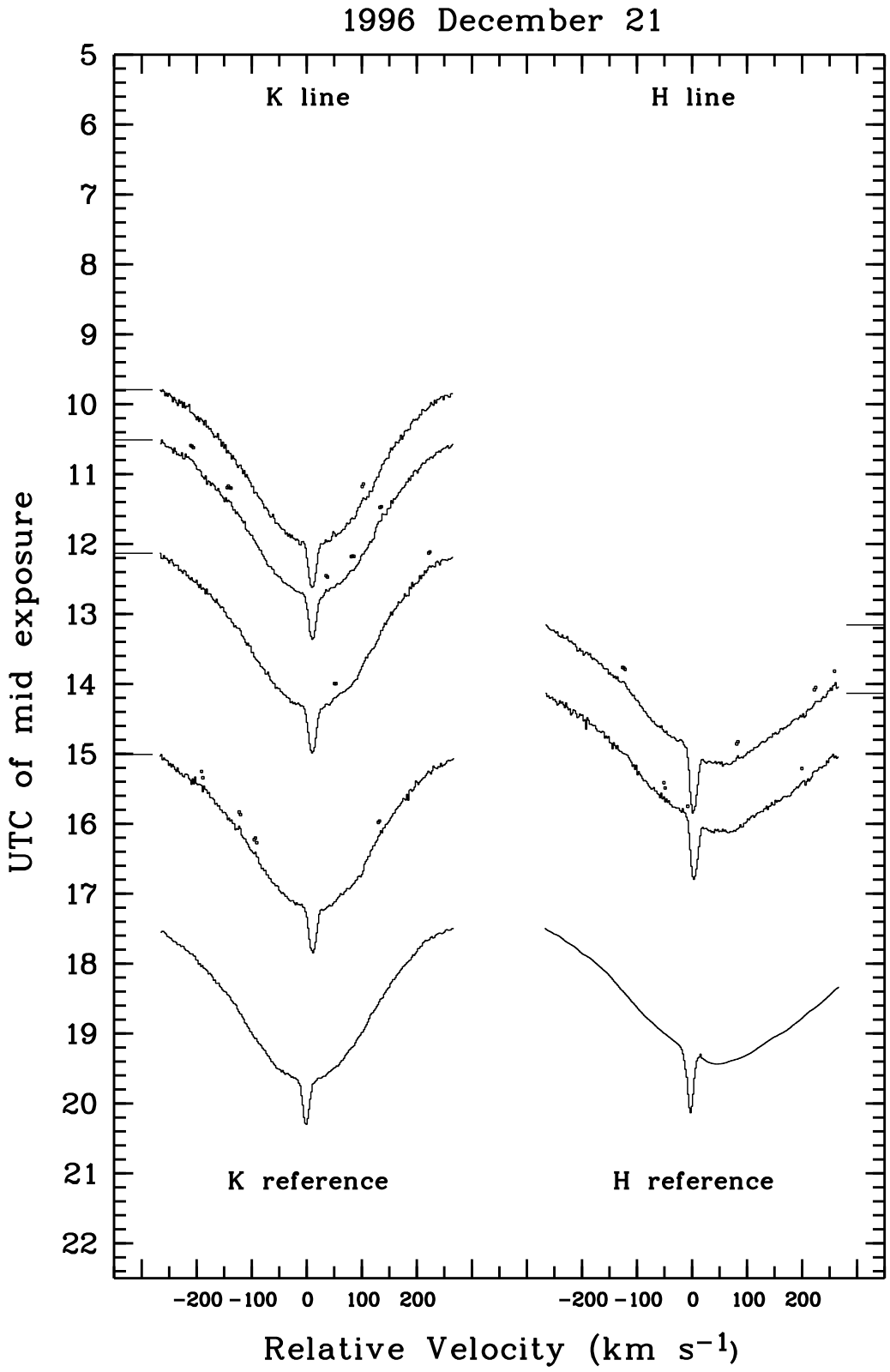

Figure 53: Observations of β Pictoris on 1996 December 21 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.